



IBGS: Image-Based Gaussian Splatting

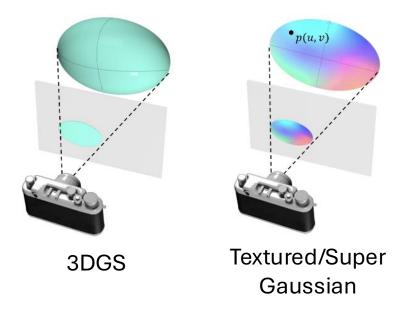
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Motivation and Contribution

- 3DGS: model Gaussian color using a low-degree spherical harmonic
 - Handle complex view-dependent color.
 - Render images with high-frequency details.
- TexturedGaussian/SuperGaussian: learns spatially varying color of each Gaussian.
 - Render images with more fine-grained details.
 - Still cannot handle complex view-specific color.
 - Increase the storage footprint significantly.

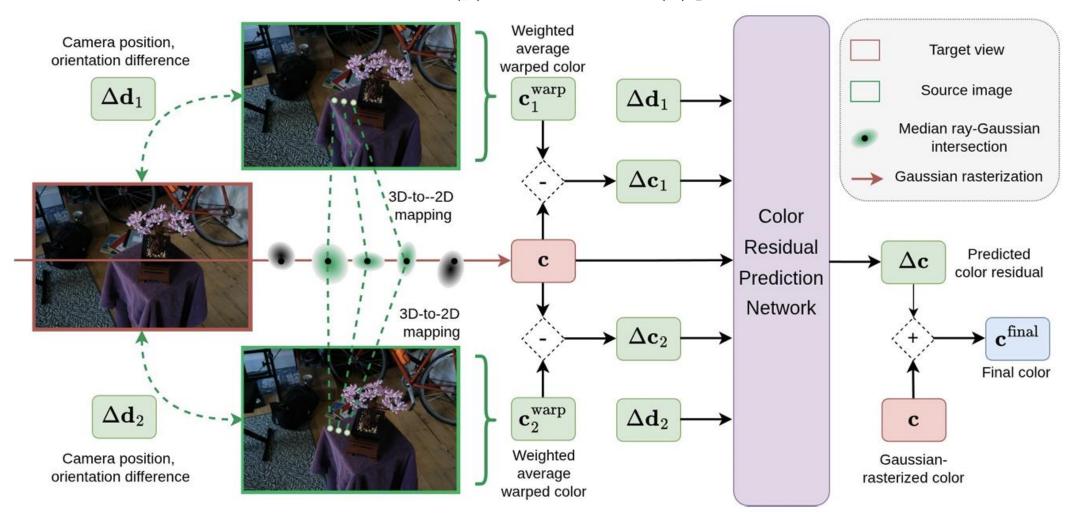


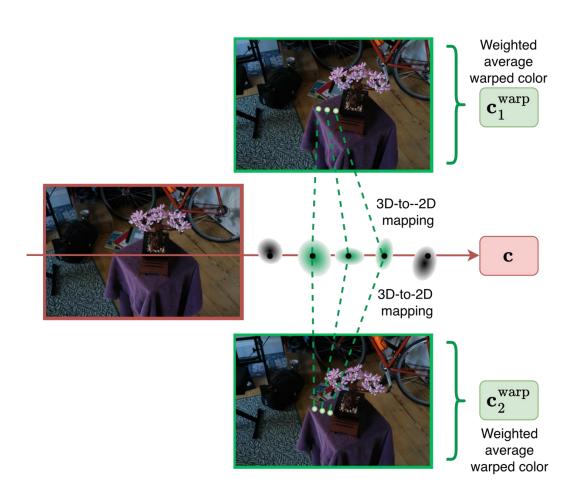
(figures obtained from SuperGaussian)

- Image-based Gaussian Splatting (IBGS):
 - Render images with fine-grained details and accurate view-specific color
 - No storage overhead.
 - Correct the exposure of an image rendered at any viewpoint.

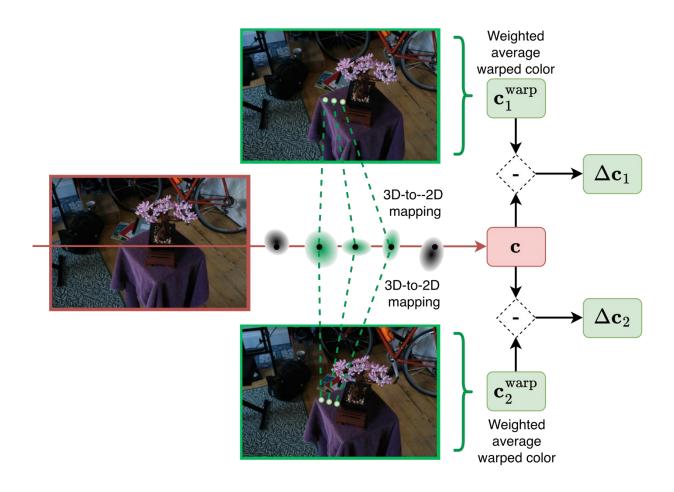
$$\mathbf{c}^{\text{final}}(\mathbf{p}) = \underbrace{\sum_{i=1}^{N} w_i \, \Psi_l(\mathbf{h}_i, \mathbf{v}_i)}_{} + \underbrace{\mathcal{F}\Big(\mathbf{c}(\mathbf{p}), \, \mathbf{d}(\mathbf{p}), \, \{\Delta \mathbf{c}_m(\mathbf{p})\}_{m=1}^{M}, \, \{\Delta \mathbf{d}_m(\mathbf{p})\}_{m=1}^{M}\Big)}_{},$$

Base color c(p) Residual $\Delta c(p)$ predicted from multi-view observations



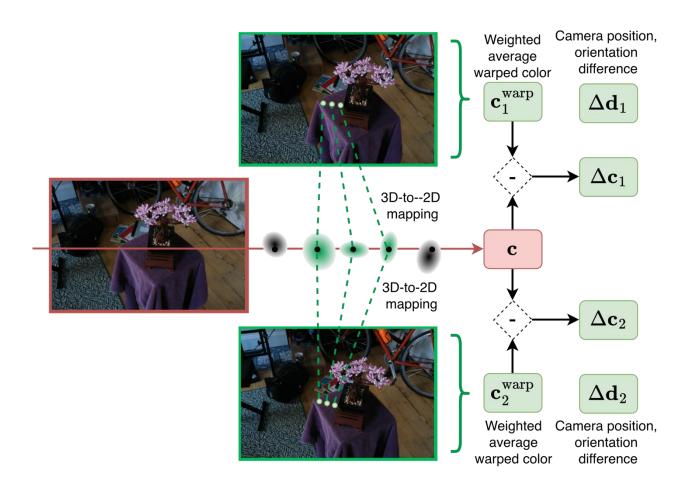


- 1. Base color **c** can is obtained via standard Gaussian rasterization.
- 2. Computing the warped colors:
 - Project camera-ray Gaussians onto multiple training images.
 - Obtain the warped colors at the projected pixels.
 - Compute the weighted average warped colors.
- \Rightarrow High-frequency details and view-dependent color information are captured in the multiview warped colors c_m^{warp} .



3. Compute the color difference:

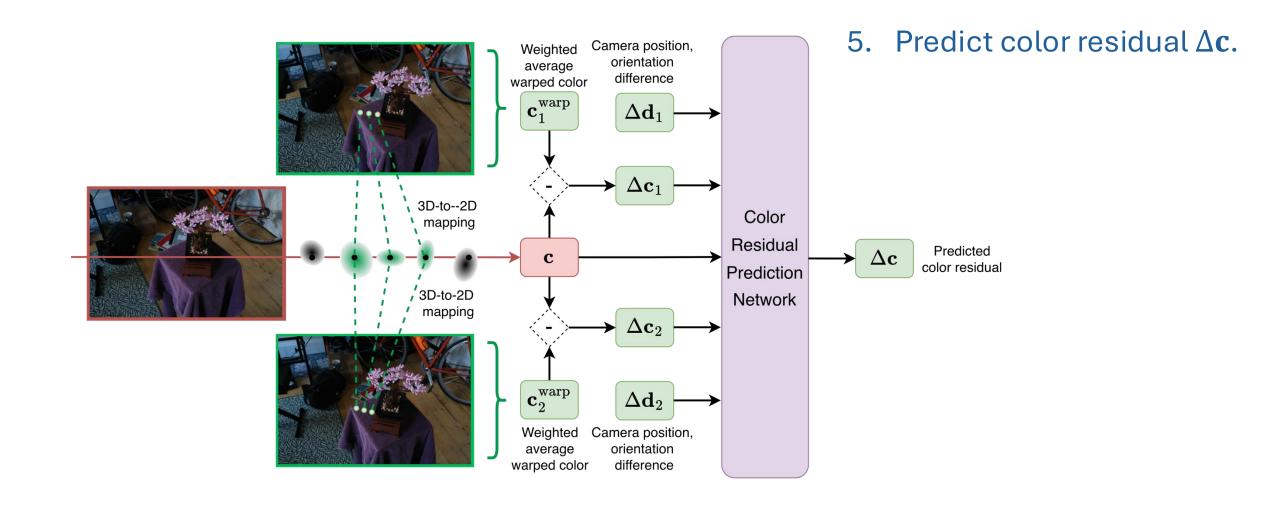
$$\Delta \mathbf{c}_m = \mathbf{c}_m^{\mathrm{warp}} - \mathbf{c}$$

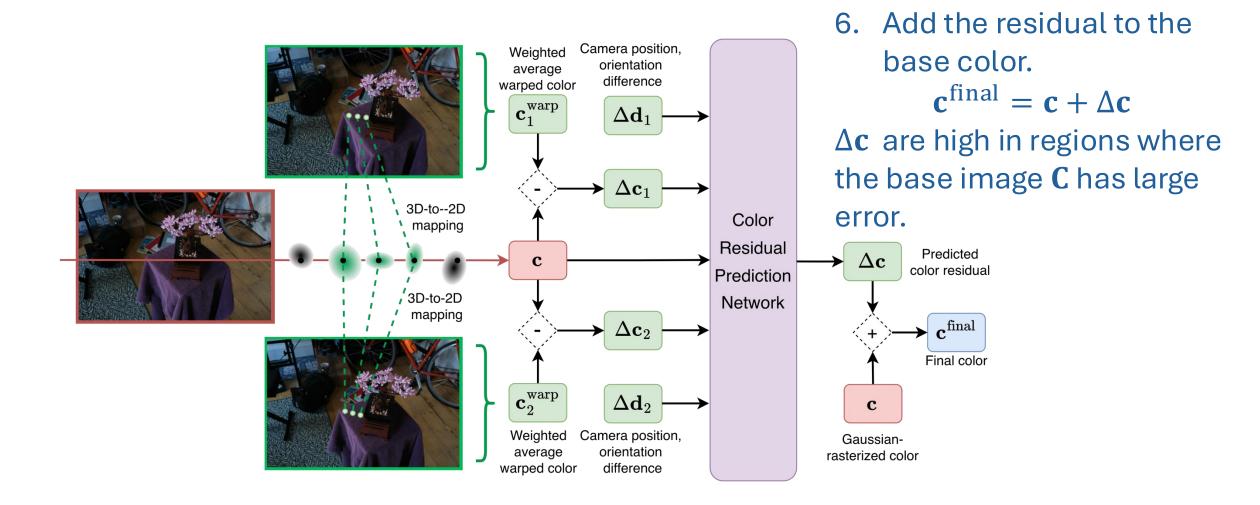


3. Compute the color difference

$$\Delta \mathbf{c}_m = \mathbf{c}_m^{\text{warp}} - \mathbf{c}$$

4. Compute camera pose difference $\Delta \mathbf{d}_m$.





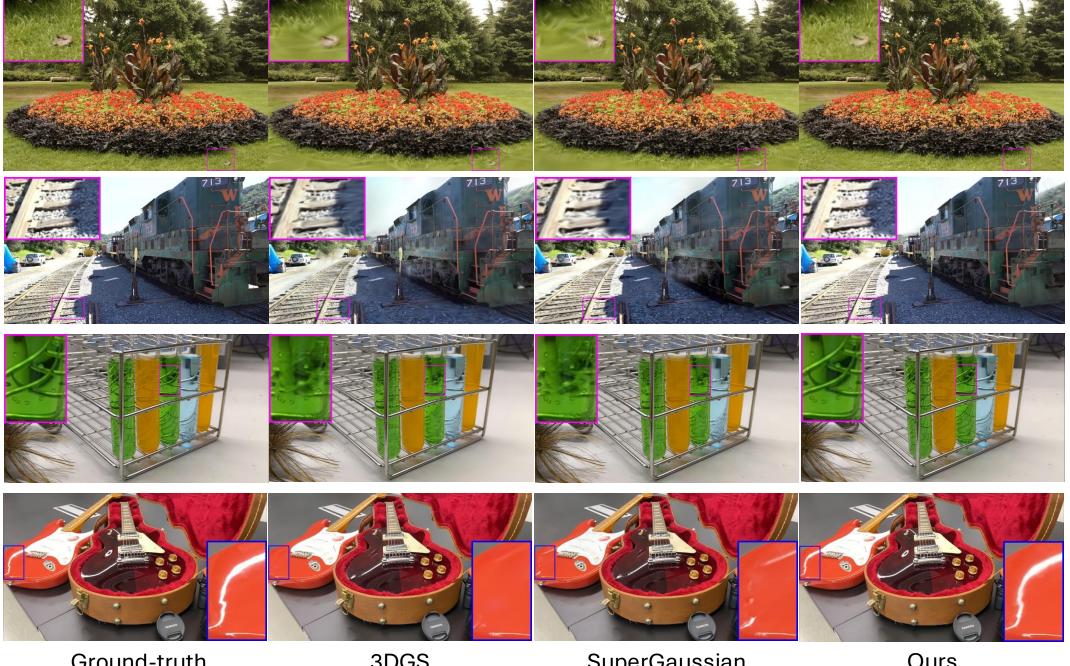
Exposure correction module

- Assume images taken at nearby locations share similar lighting condition.
- ⇒ Correct the exposure of a rendered image by mimicking that of its nearest training view.
- Find a matrix representing the target view's exposure:

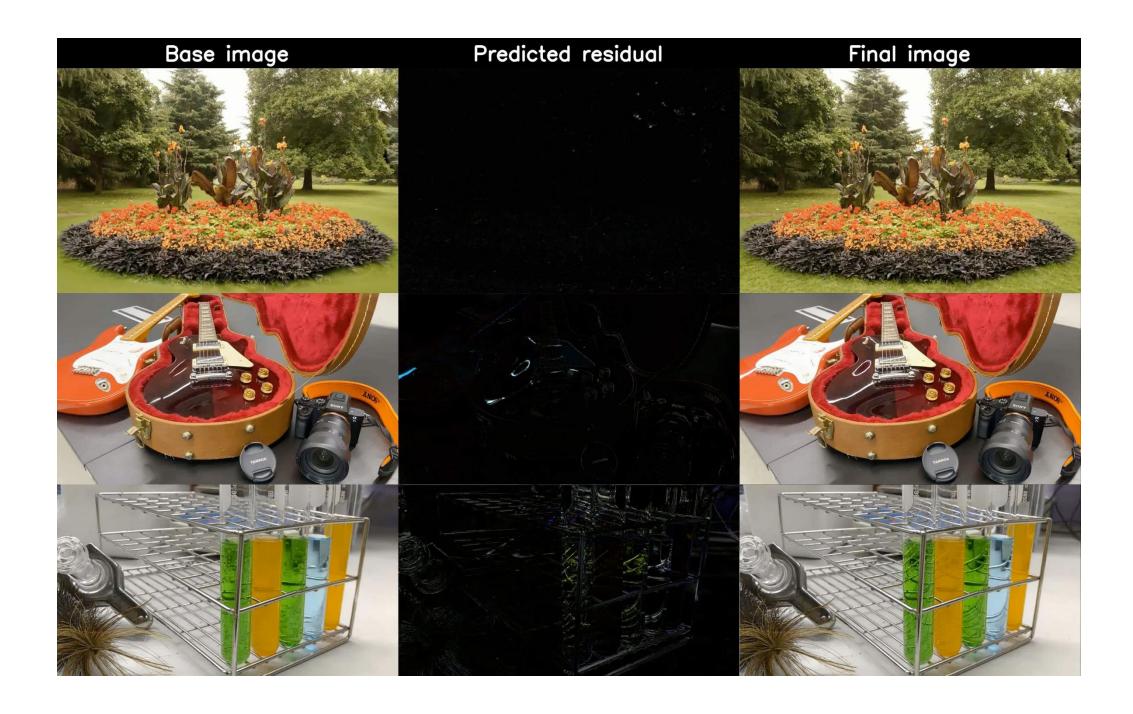
$$\mathbf{A}^* = \arg\min_{\mathbf{A} \in \mathbb{R}^{3 \times 4}} \sum_{p \in \chi} \left| \left| \mathbf{A} \begin{bmatrix} \mathbf{C}(\mathbf{p}) \\ 1 \end{bmatrix} - \mathbf{C}^{\text{near}}(\mathbf{p}) \right| \right|_2^2$$

• Apply A^* to correct the exposure of the rendered image:

$$\mathbf{C}^{\mathrm{corr}}(\mathbf{p}) = \mathbf{A}^* \begin{bmatrix} \mathbf{C}(\mathbf{p}) \\ 1 \end{bmatrix}$$



3DGS SuperGaussian Ground-truth Ours



Thanks for watching

Let's have a chat if you find this work interesting!

Morning session, Thursday Dec 4th
Poster Session 3



Project page